

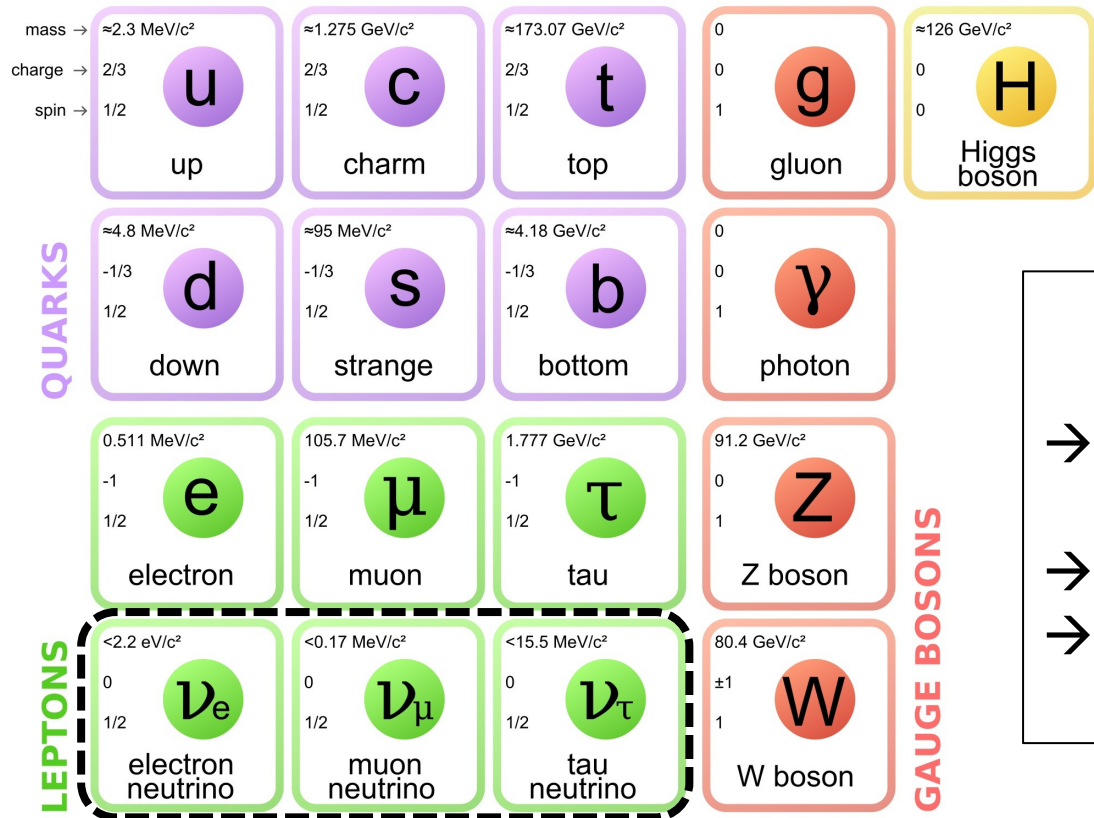


Neutrino oscillations and Earth tomography with the ORCA deep-sea detector

Simon Bourret

Journée des doctorants APC, 10 November 2016

Neutrino oscillations



NEUTRINO

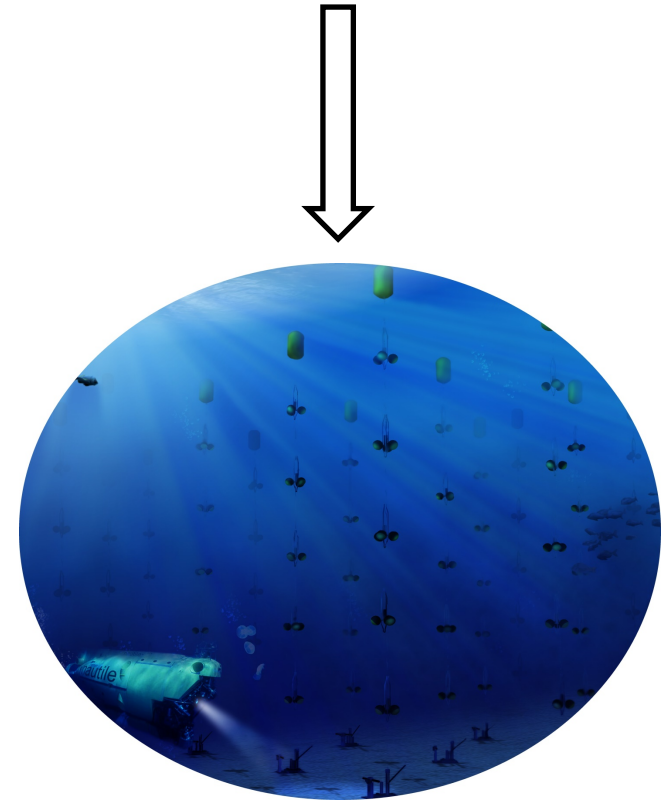
- interacts very weakly with matter
- massless in Standard Model
- actually not massless... but almost !

- ✓ 3 generations or flavours
- ✓ flavour mixing

(Flavour mixing) & (Distinct masses)
 ↓
Flavour oscillations

ORCA ?

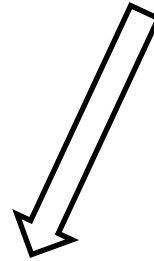
« Oscillation Research with Cosmics in the Abyss »



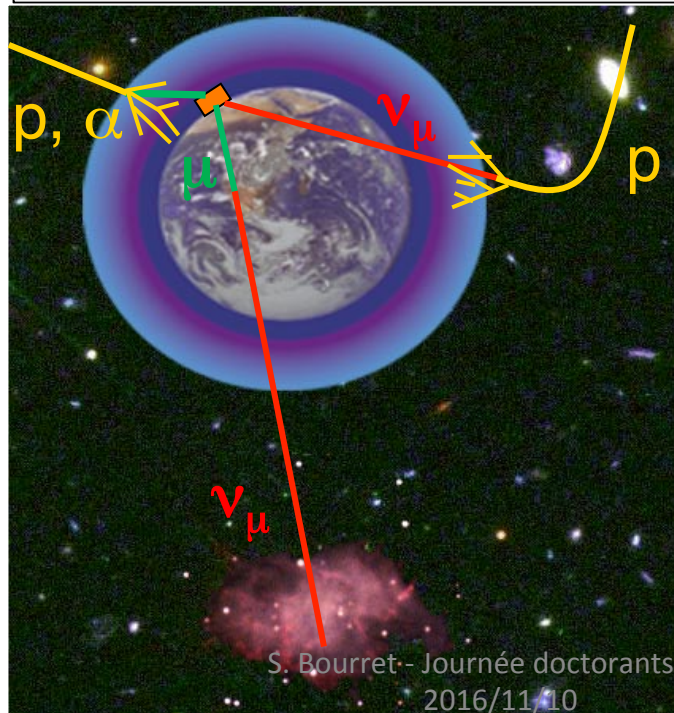
- ✓ *Very large detection volume*
- ✓ *Background reduction*

ORCA ?

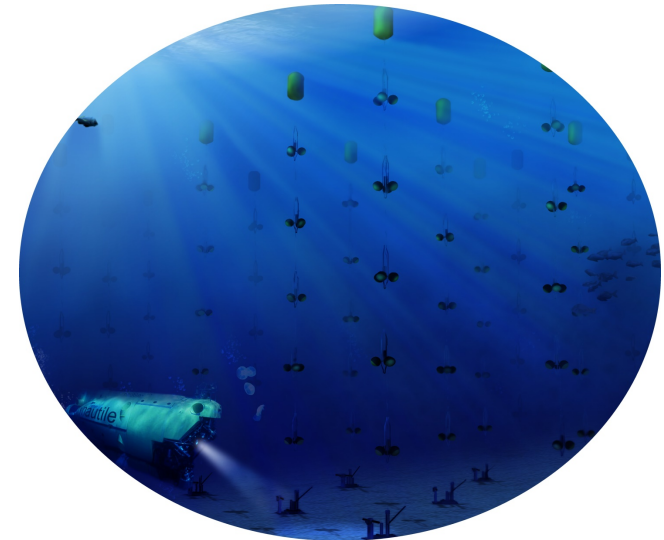
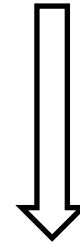
« Oscillation Research with Cosmics in the Abyss »



Cosmic rays induce atmospheric neutrinos



S. Bourret - Journée doctorants APC
2016/11/10



- ✓ *Very large detection volume*
- ✓ *Background reduction*

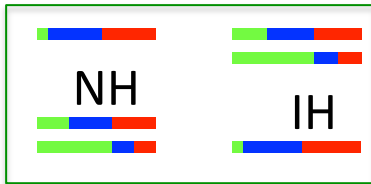
ORCA ?

« Oscillation Research with Cosmics in the Abyss »



Neutrino flavour oscillations

→ Neutrino mass hierarchy



→ Earth tomography

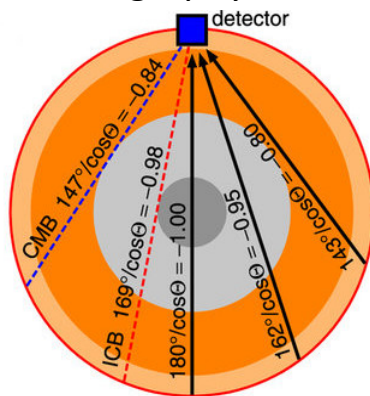
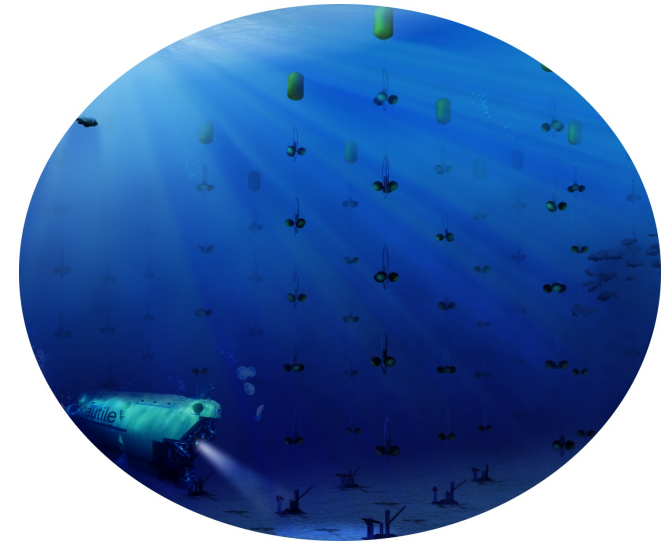


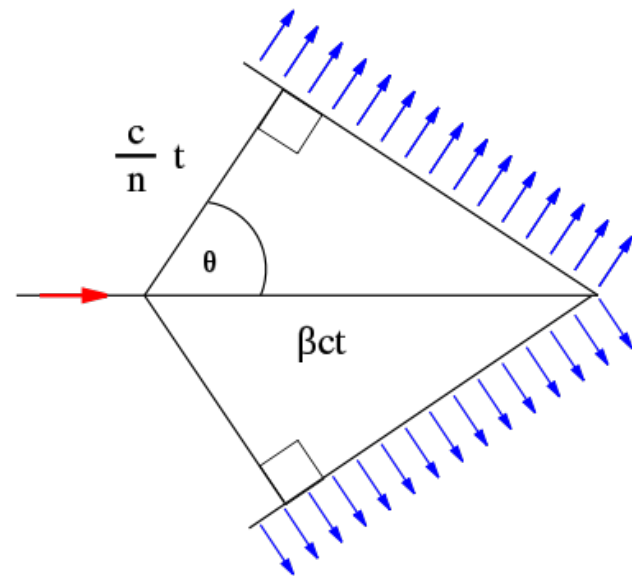
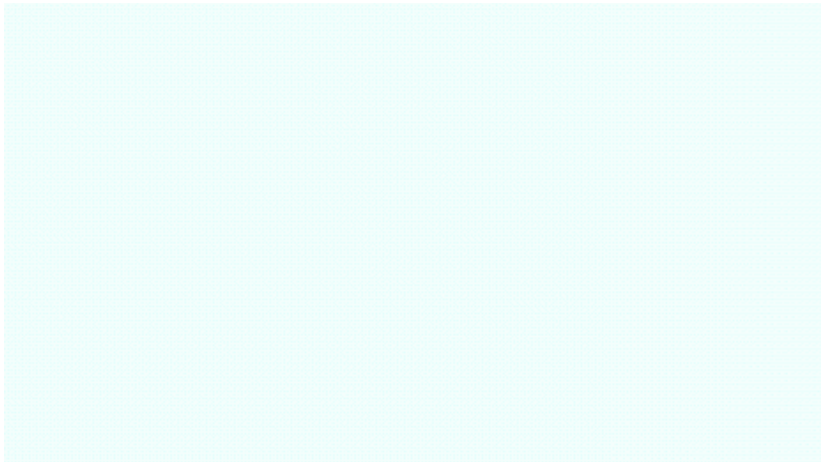
figure C. Rott

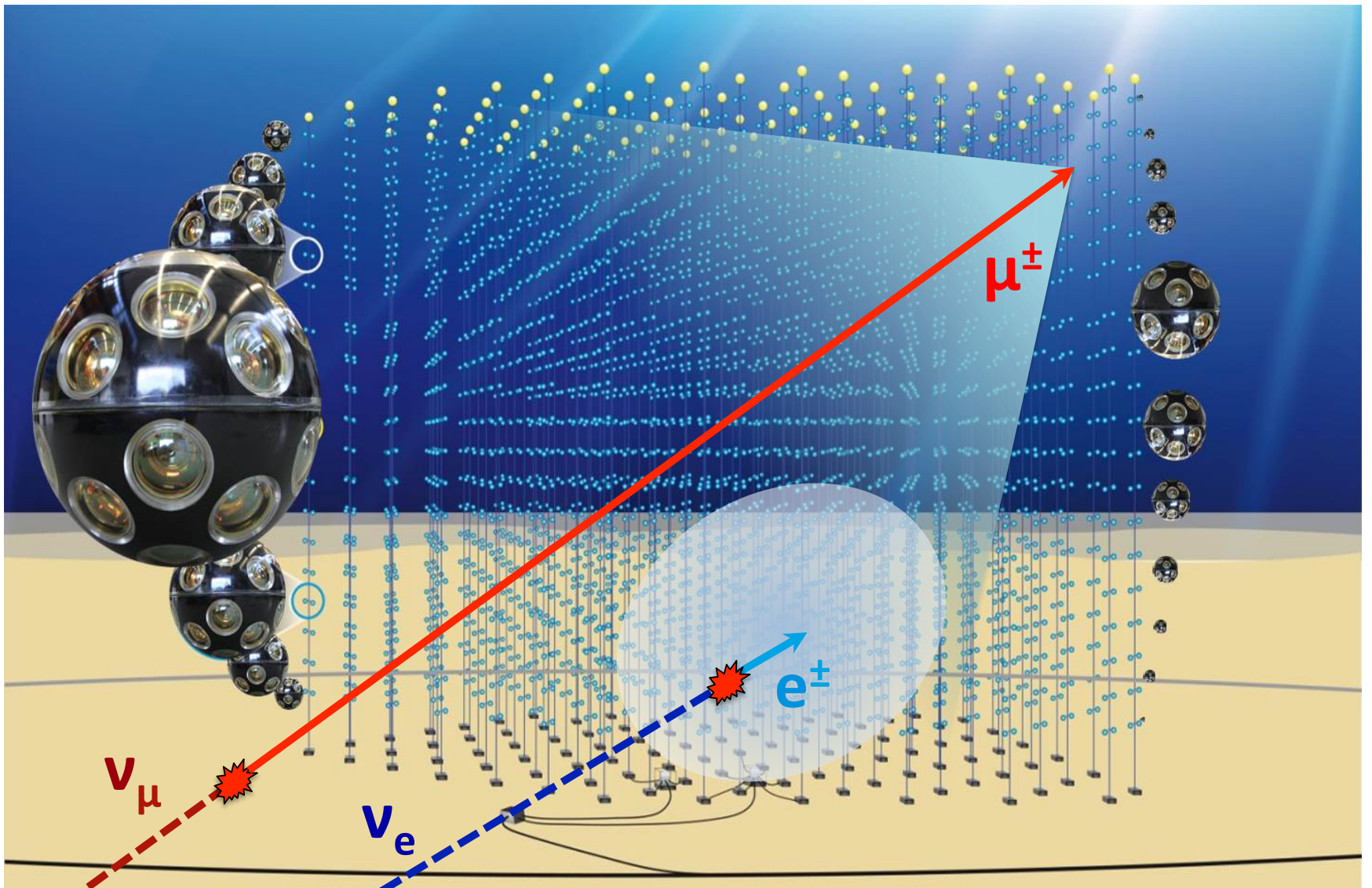


- ✓ *Very large detection volume*
- ✓ *Background reduction*

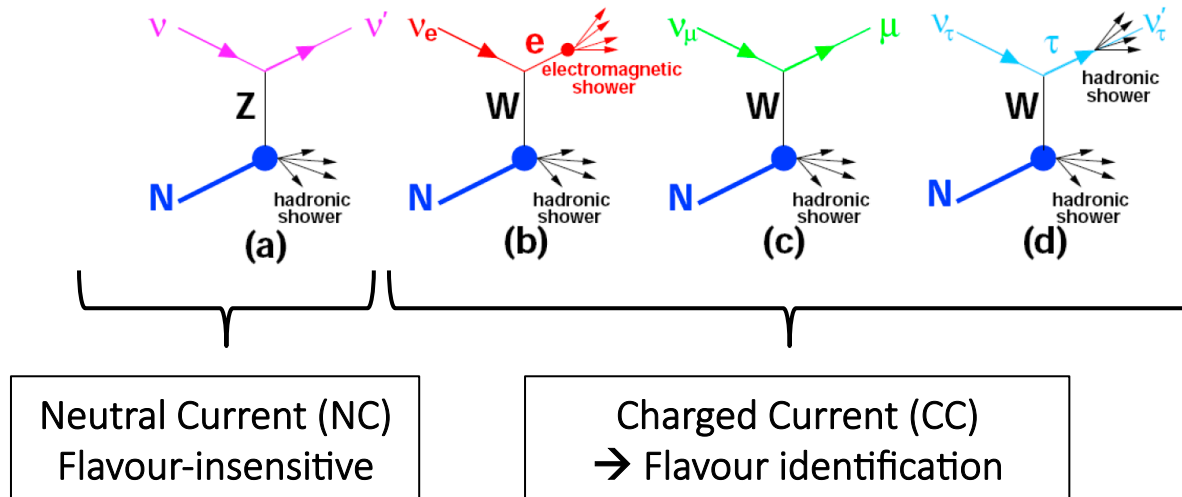
Detection principle : Cherenkov radiation

- Particles moving faster than the phase velocity of light in a dielectric medium produce coherent radiation → « Cherenkov cone »
- Neutrino-nucleus scattering → Charged secondary particles → Cherenkov emission
- Cherenkov angle $\theta \sim 43^\circ$, energy-independent for highly relativistic particles in water

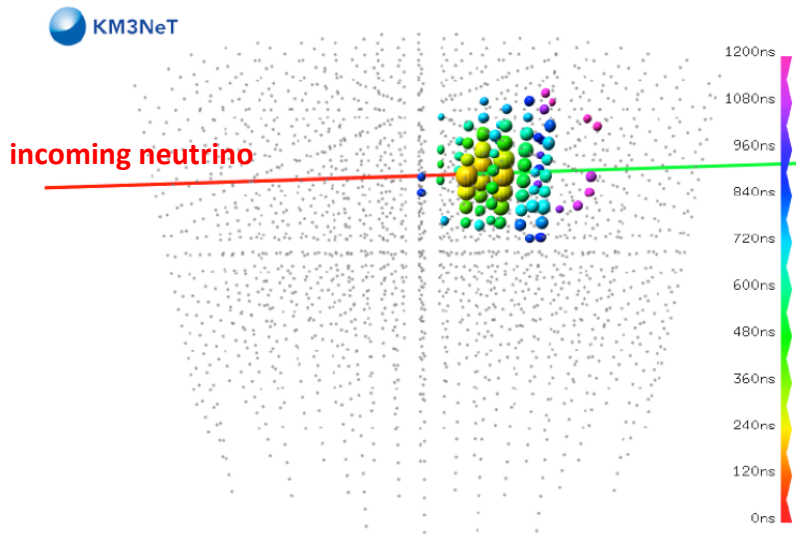
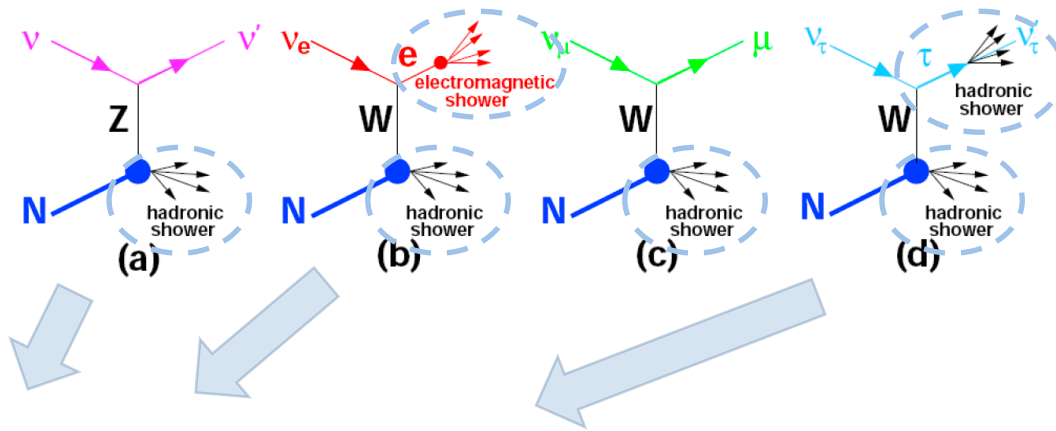




Neutrino interactions and signatures

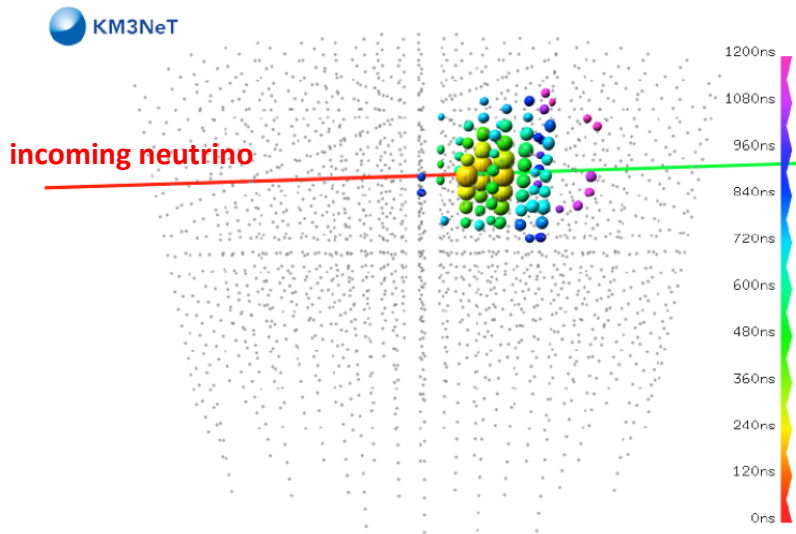
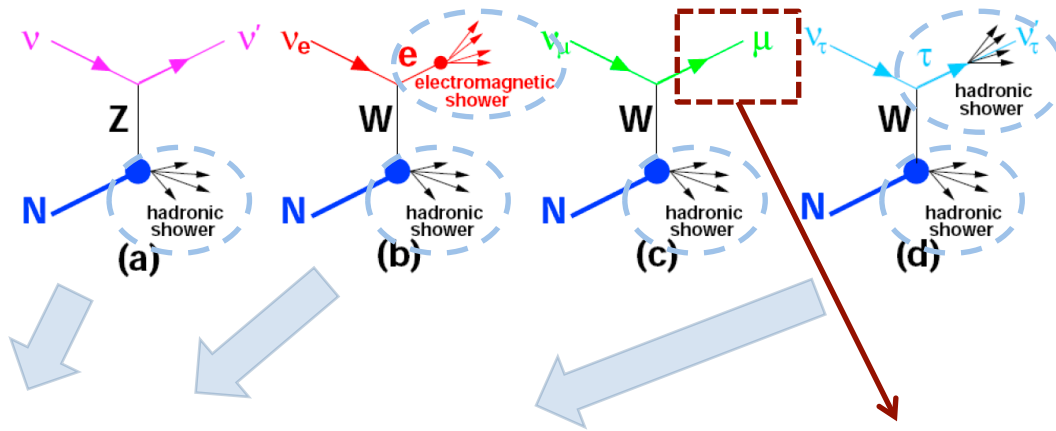


Neutrino interactions and signatures

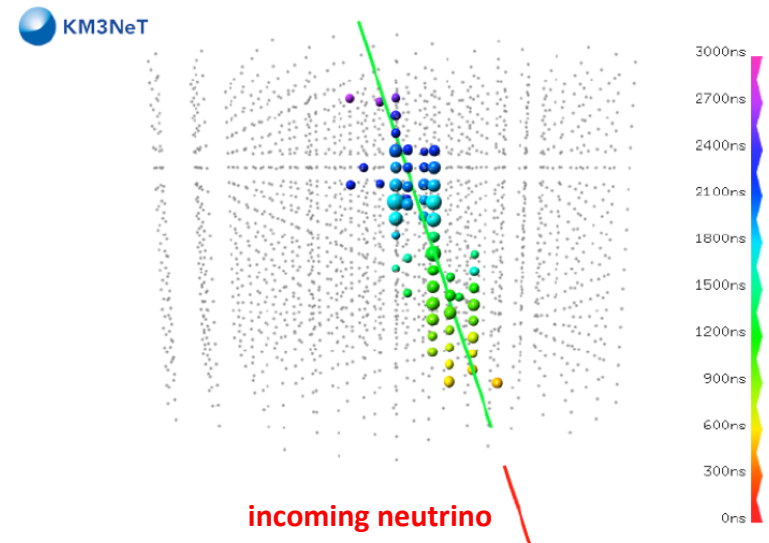


Cascade-like event

Neutrino interactions and signatures



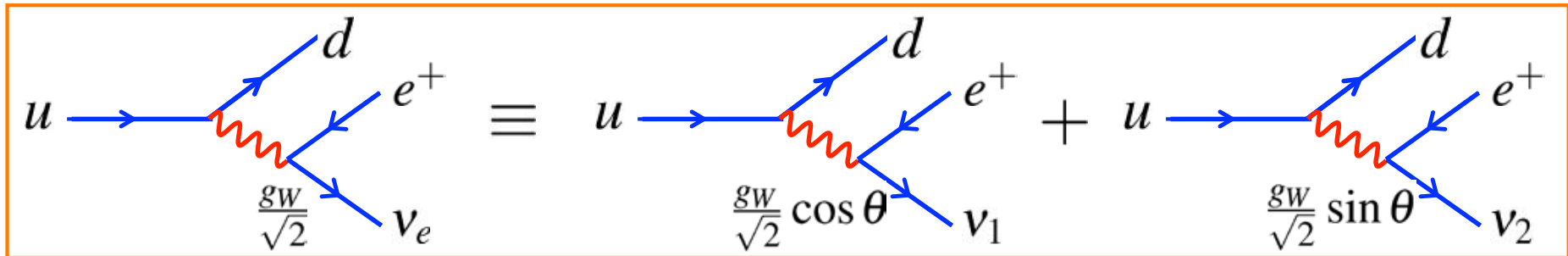
Cascade-like event



Track-like event

Weak and mass eigenstates

- Let's assume two massive neutrinos.
- Neutrinos are produced and interact as weak eigenstates ν_e, ν_μ
- The weak eigenstates are coherent superposition of the fundamental mass eigenstates ν_1, ν_2
- The mass eigenstates are the solutions of the free Hamiltonian and represent the propagation of the neutrinos in space.



$\theta =$ mixing angle

$$\begin{array}{ccc}
 \text{Mass eigenstates} & \begin{array}{c} \nearrow \\ \left(\begin{array}{c} \nu_1 \\ \nu_2 \end{array} \right) = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} \\
 \searrow \end{array} & & \begin{array}{c} \nwarrow \\ \text{Weak eigenstates} \end{array}
 \end{array}$$

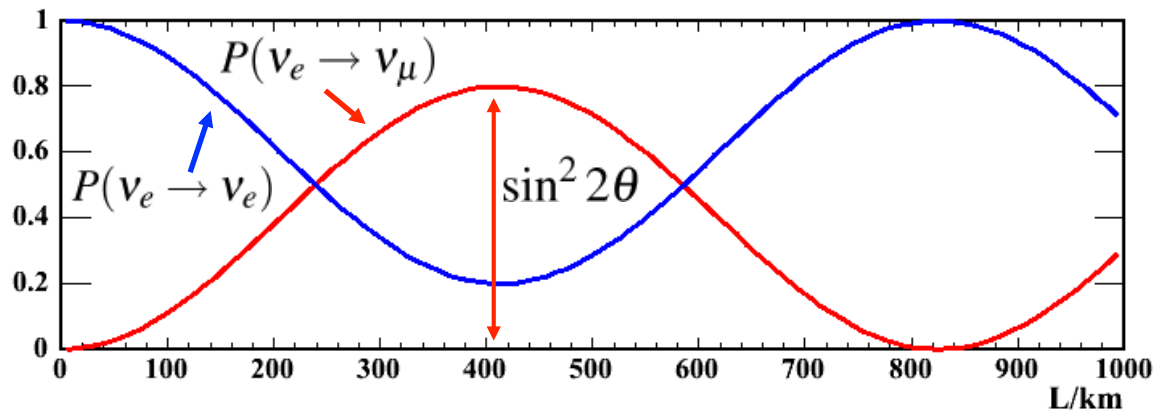
Two neutrino flavor oscillations

- If the masses are different, the two mass eigenstates propagate with different phases. The resulting state is the quantum mechanical superposition.
- At a distance L from the source, neutrinos will be detected in a given flavor (=weak eigenstate) with a probability depending the neutrino energy:

$$P_{\nu_e \rightarrow \nu_\mu} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)$$

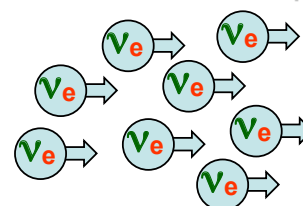
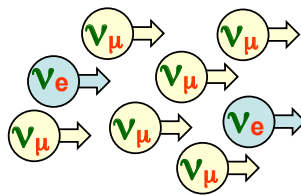
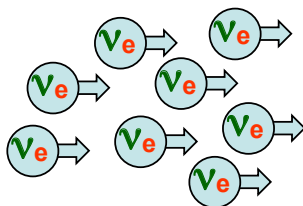
Distance source-detector: L
 Neutrino energy: E
 $\Delta m^2 \equiv m_2^2 - m_1^2$

•e.g. $\Delta m^2 = 0.003 \text{ eV}^2$, $\sin^2 2\theta = 0.8$, $E_\nu = 1 \text{ GeV}$



•wavelength

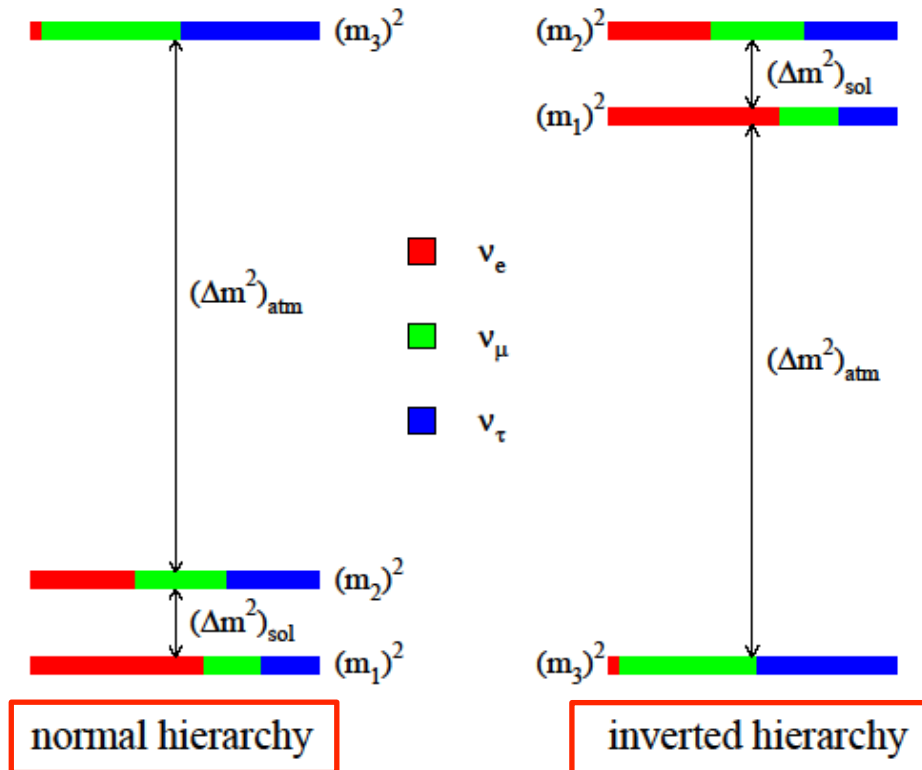
$$\lambda_{\text{osc}} = \frac{4\pi E}{\Delta m^2}$$



(Two neutrinos flavor)

Neutrino Mass Hierarchy (NMH)

- Standard oscillation experiments are not sensitive to the sign of Δm^2_{ij}
- The sign of $\Delta m^2_{21} = (\Delta m^2)_{\text{sol}}$ has been determined from solar oscillations
- Two situations : normal hierarchy (NH) vs. inverted hierarchy (IH)

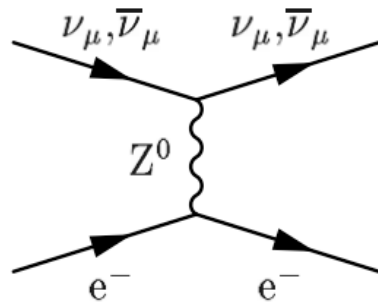


Matter effects in the Earth (MSW effect)

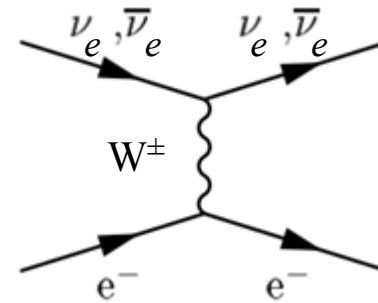
Propagation of atmospheric neutrinos through the Earth

→ Scattering off nucleons and electrons (« coherent forward scattering »)

- neutrino-nucleon : all neutrino flavours involved
- neutrino-electron :



Neutral current : all flavours



Charged current : ν_e only

→ Electronic and muonic components of the atmospheric neutrino flux do not see the same matter potential

Matter effects: simple picture

- Vacuum oscillations in 2-flavour approximation

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2(2\theta) \sin^2\left(1.27 \times \Delta m^2 \times \frac{L}{E}\right)$$

- 2-flavour oscillation in constant matter density with electron density N_e

$$\theta \rightarrow \theta_m$$

$$\tan(2\theta_m) = \frac{\Delta m^2 \sin^2 2\theta}{-A + \Delta m^2 \cos 2\theta}$$

Matter potential
⇒ mass eigenstates shifted
⇒ effective mixing angle

$$A(\nu) = +2\sqrt{2}G_F \times E \times N_e$$

$$A(\bar{\nu}) = -2\sqrt{2}G_F \times E \times N_e$$

Resonance when $A = A_{\text{res}}$

Matter effects: less simple picture

$$\begin{aligned}
 P_{3\nu}^m(\nu_\mu \rightarrow \nu_\mu) \approx & 1 - \sin^2 2\theta_{23} \cos^2 \theta_{13}^m \sin^2 \left(\frac{(\Delta m_{31}^2 + \Delta^m m^2)L}{8E_\nu} + \frac{AL}{4} \right) \\
 & - \sin^2 2\theta_{23} \sin^2 \theta_{13}^m \sin^2 \left(\frac{(\Delta m_{31}^2 - \Delta^m m^2)L}{8E_\nu} + \frac{AL}{4} \right) \\
 & - \sin^4 \theta_{23} \sin^2 2\theta_{13}^m \sin^2 \left(\frac{\Delta^m m^2 L}{4E_\nu} \right)
 \end{aligned}$$

$$P_{3\nu}^m(\nu_\mu \rightarrow \nu_e) \approx \sin^2 \theta_{23} \sin^2 2\theta_{13}^m \sin^2 \left(\frac{\Delta^m m^2 L}{4E_\nu} \right)$$

$$\sin^2 2\theta_{13}^m \equiv \sin^2 2\theta_{13} \left(\frac{\Delta m_{31}^2}{\Delta^m m^2} \right)^2$$

$$\Delta^m m^2 \equiv \sqrt{(\Delta m_{31}^2 \cos 2\theta_{13} - 2E_\nu A)^2 + (\Delta m_{31}^2 \sin 2\theta_{13})^2}$$

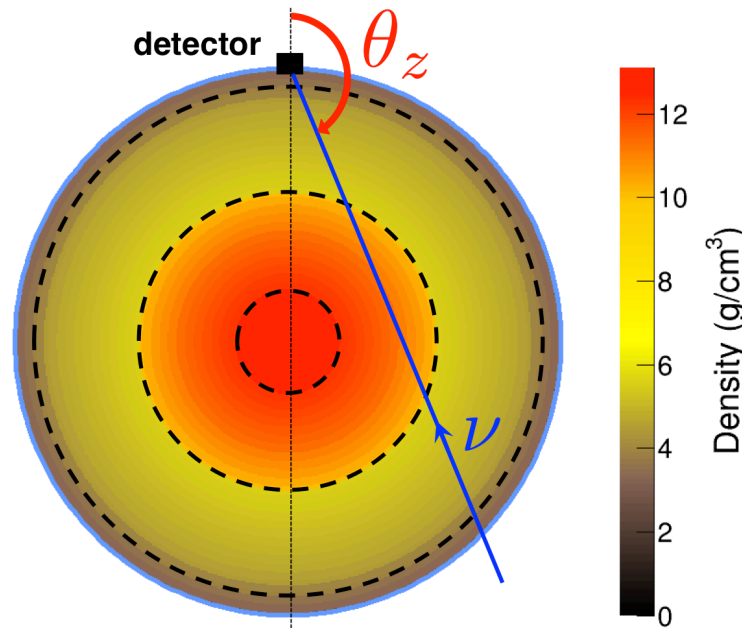
$$A = \pm \sqrt{2} G_F N_e$$

Matter effects: less simple picture

Varying matter density in the Earth

→ Use discretized model with constant density steps based on geophysics standard model

→ Step-by-step numerical resolution of evolution equation



$$L = 2R_{\text{Earth}} \cos \theta_z$$

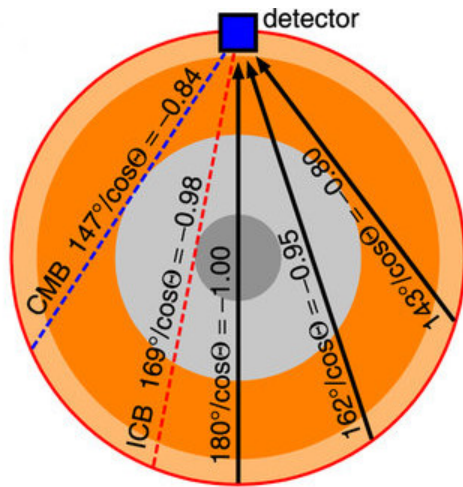
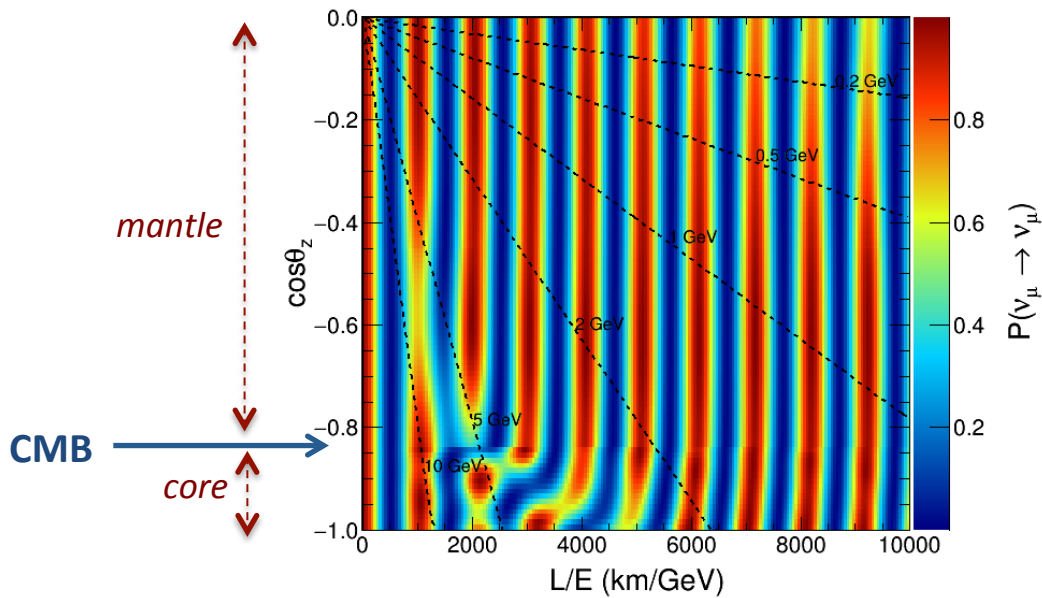


figure C. Rott



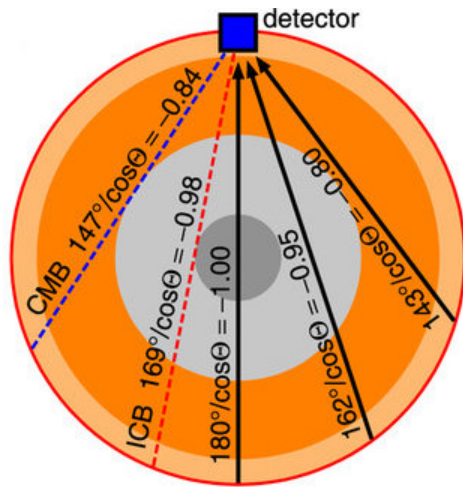
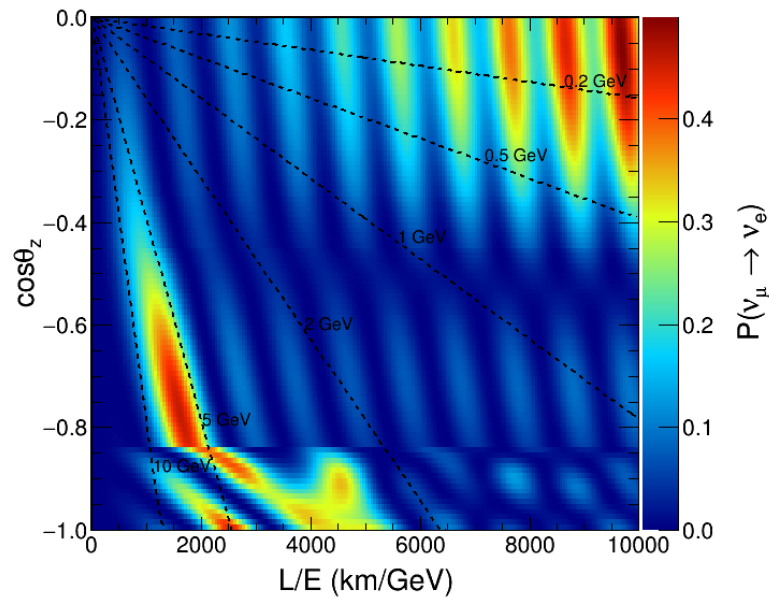
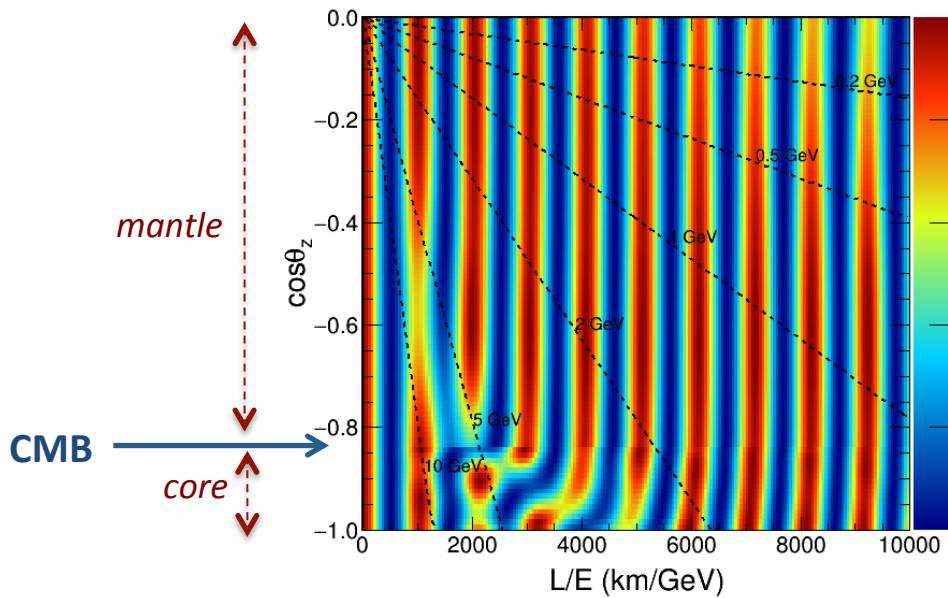


figure C. Rott



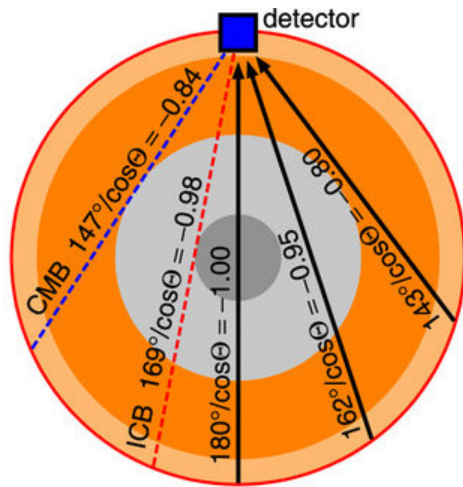
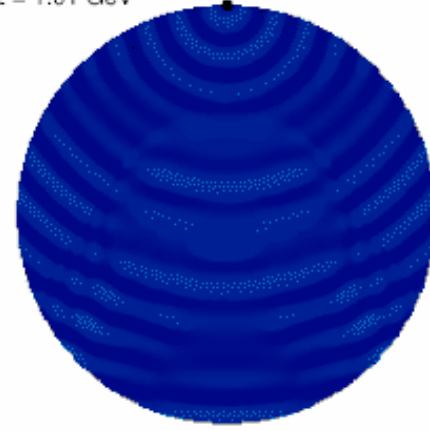
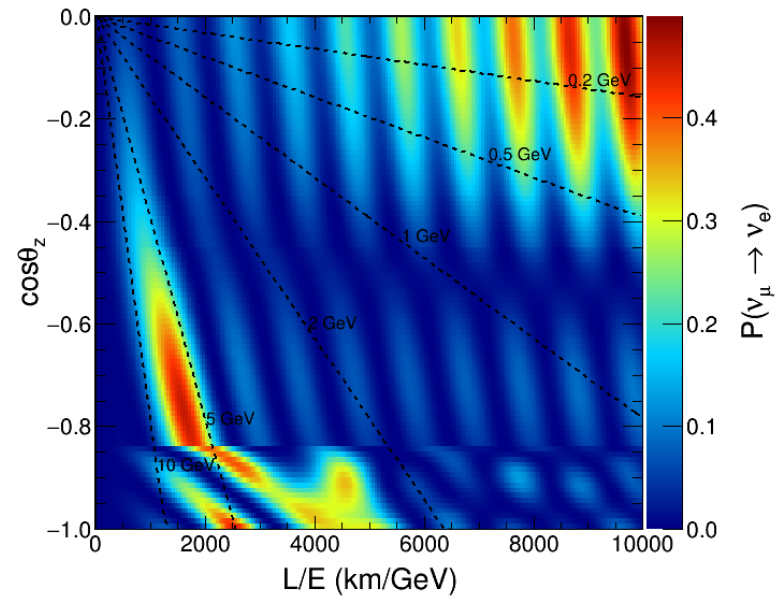
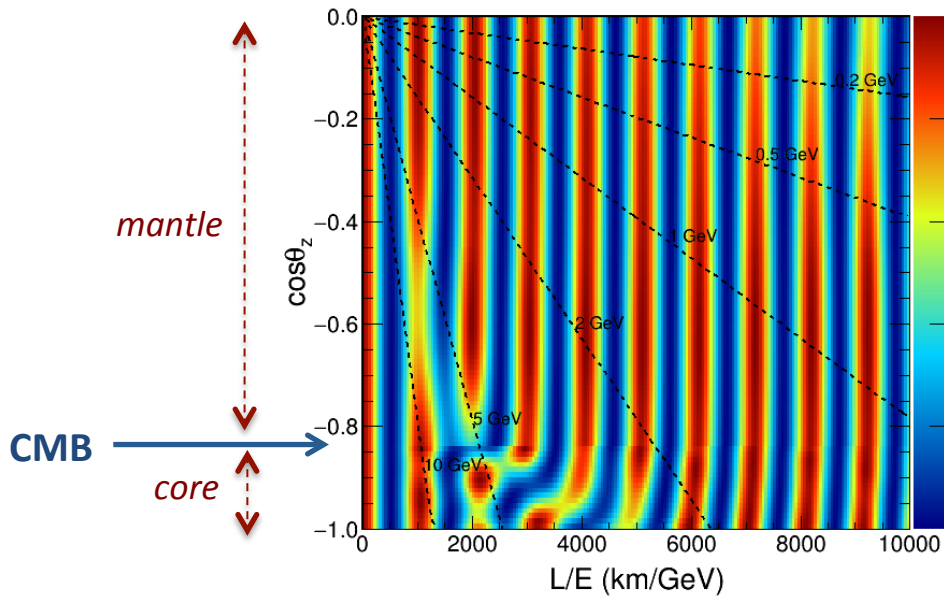


figure C. Rott

E = 1.01 GeV



animation J. Coelho



Electron density vs. matter density

Electron to mass density ratio \rightarrow depends on the **chemical composition** of the medium

w_i = weight fraction of chemical element i

Proton-to-nucleon ratio ≈ 0.5

$$\frac{Z}{A} \equiv \sum w_i \frac{Z_i}{A_i} \quad \longrightarrow \quad n_e = \frac{N_A}{m_n} \times \frac{Z}{A} \times \rho_{\text{matter}}$$

Electron density
 \Rightarrow measured with
neutrino oscillations

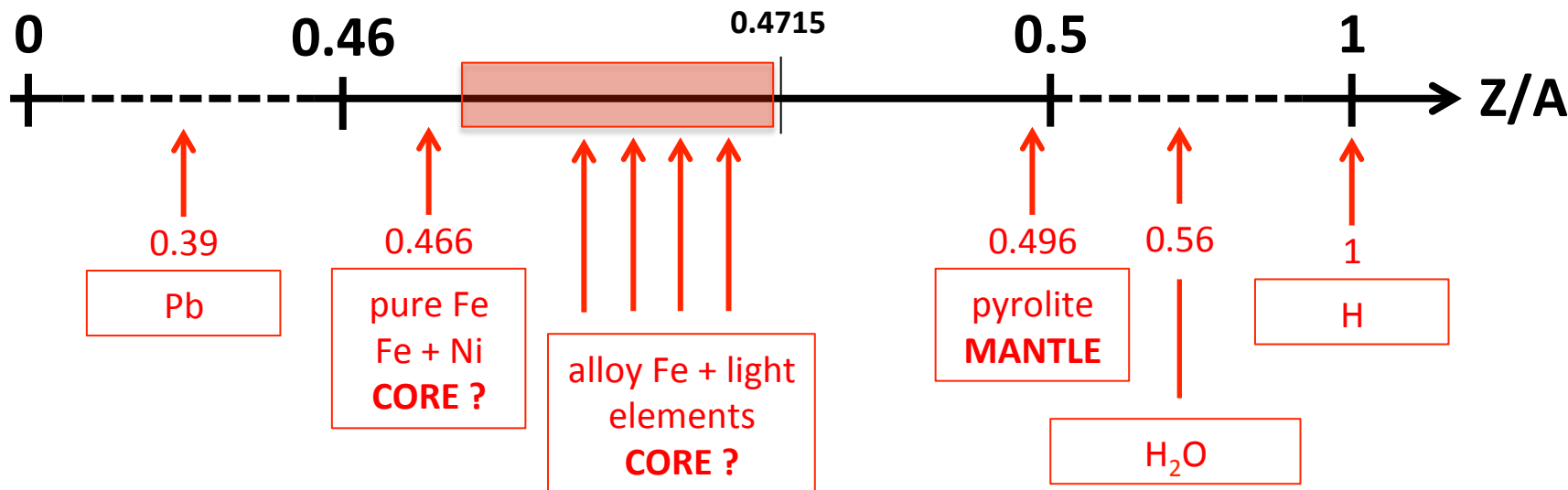
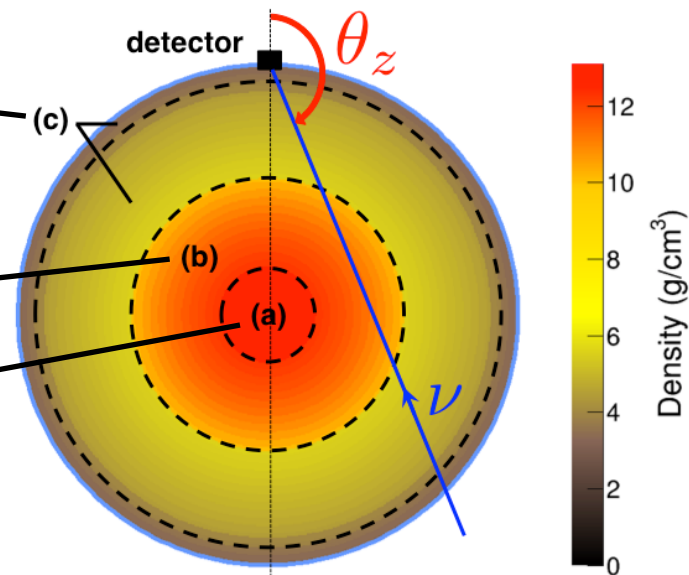
Matter density (\rightarrow nucleons)
 \Rightarrow measured with geophys.
techniques

Earth composition: knowns and unknowns

Mantle and crust: $R_{\text{ext}} \approx 6400$ km
silicate rocks: Si, O, Mg, Fe, Na, K, ...

Outer core: $R_{\text{ext}} \approx 3480$ km
liquid Fe + few % Ni + **<1% light elements ?**

Inner core: $R_{\text{ext}} \approx 1220$ km
solid Fe

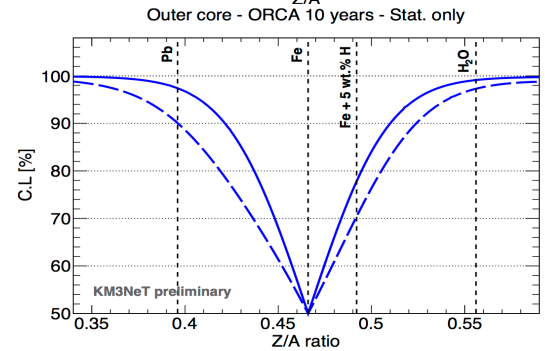
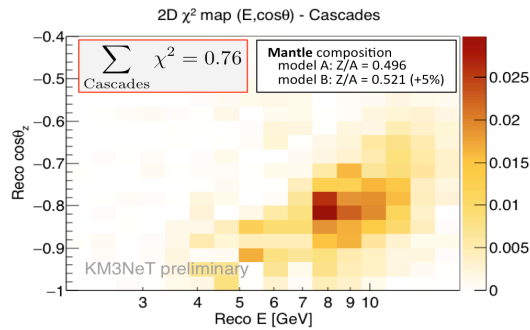
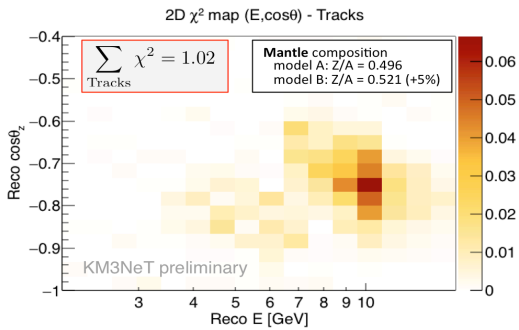
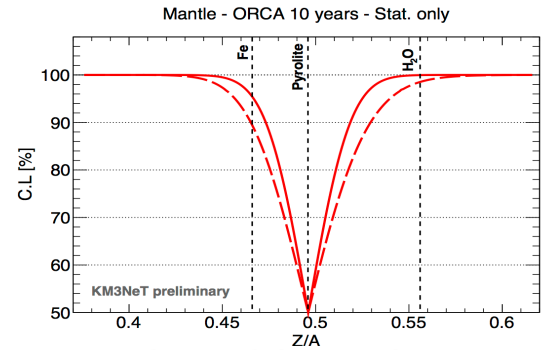
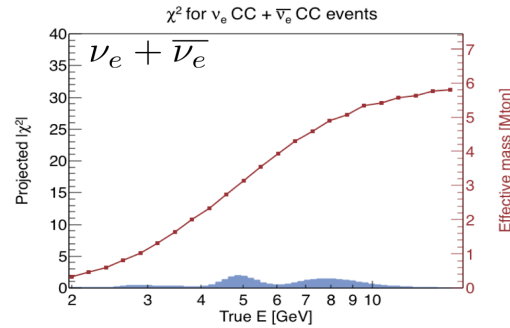
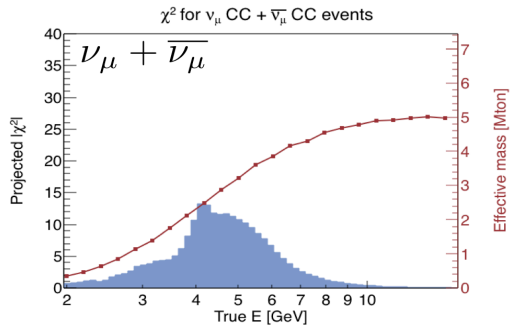
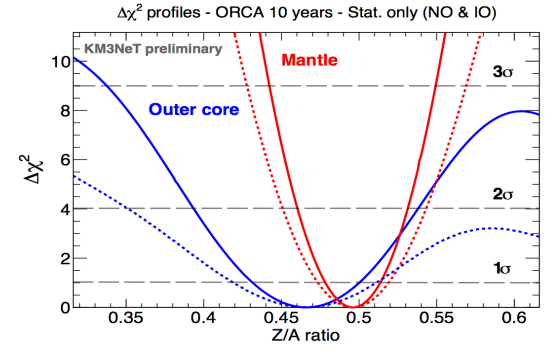
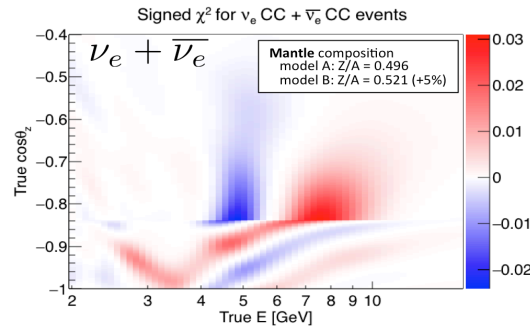
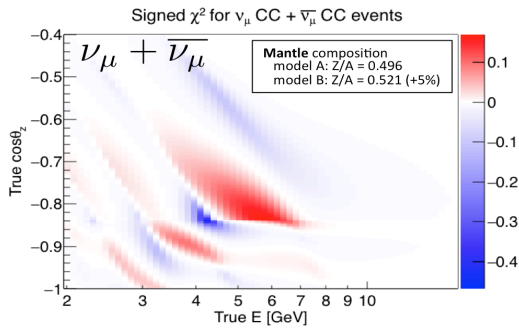


22

From the principle to the measurement

- Limitations:
 - poor detection efficiency below 2 GeV (\approx max at 10 GeV)
 - limited angular and energy resolution at low energy
 - no neutrino / antineutrino distinction on event-by-event basis
 - flavour ID limited to 2 broad classes: track and showers
- Detector response based on extensive Monte Carlo simulations
- Sensitivity evaluation: statistical model taking into account
 - statistical fluctuation
 - systematic uncertainties due to unknown parameters in the model: oscillation physics, detector response parameters, etc.

Results → come to the PHD seminar



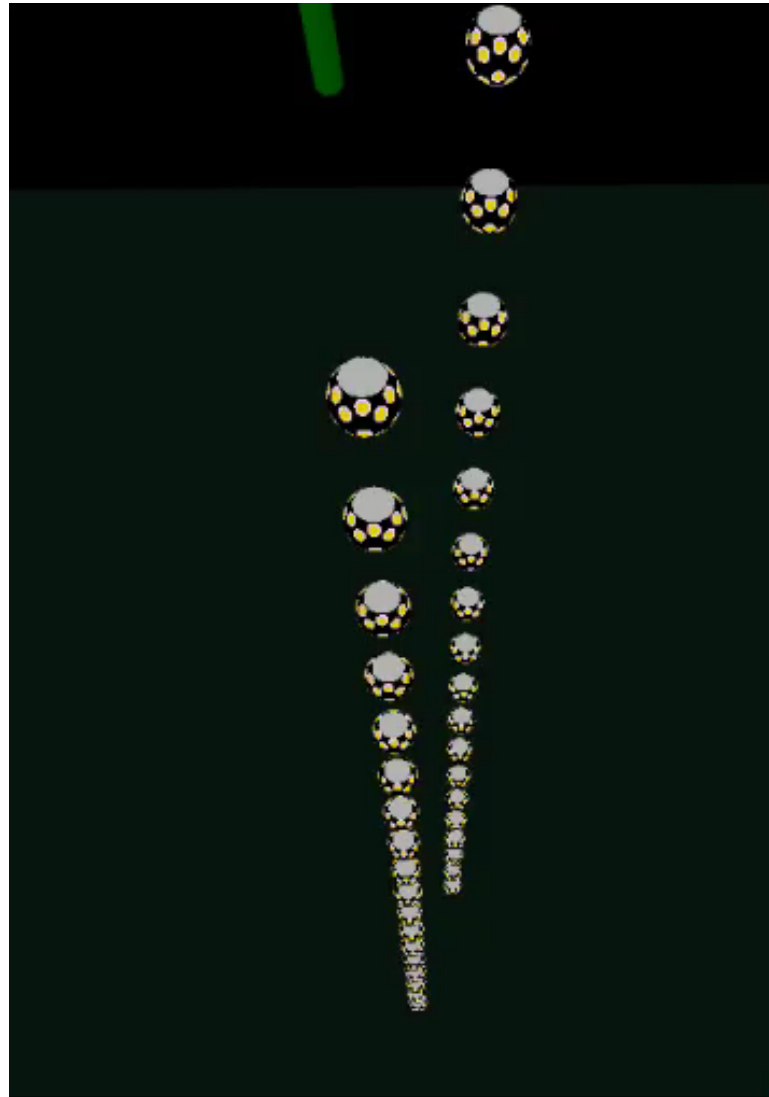
Thank you for your attention !

& come to the PHD seminar **next Thursday**

Thursday 17 Nov. 14h30

(once every 2 weeks, tell us if you are not on the mailing list)

Bonus: first KM3NeT 2-strings data



Additional slides

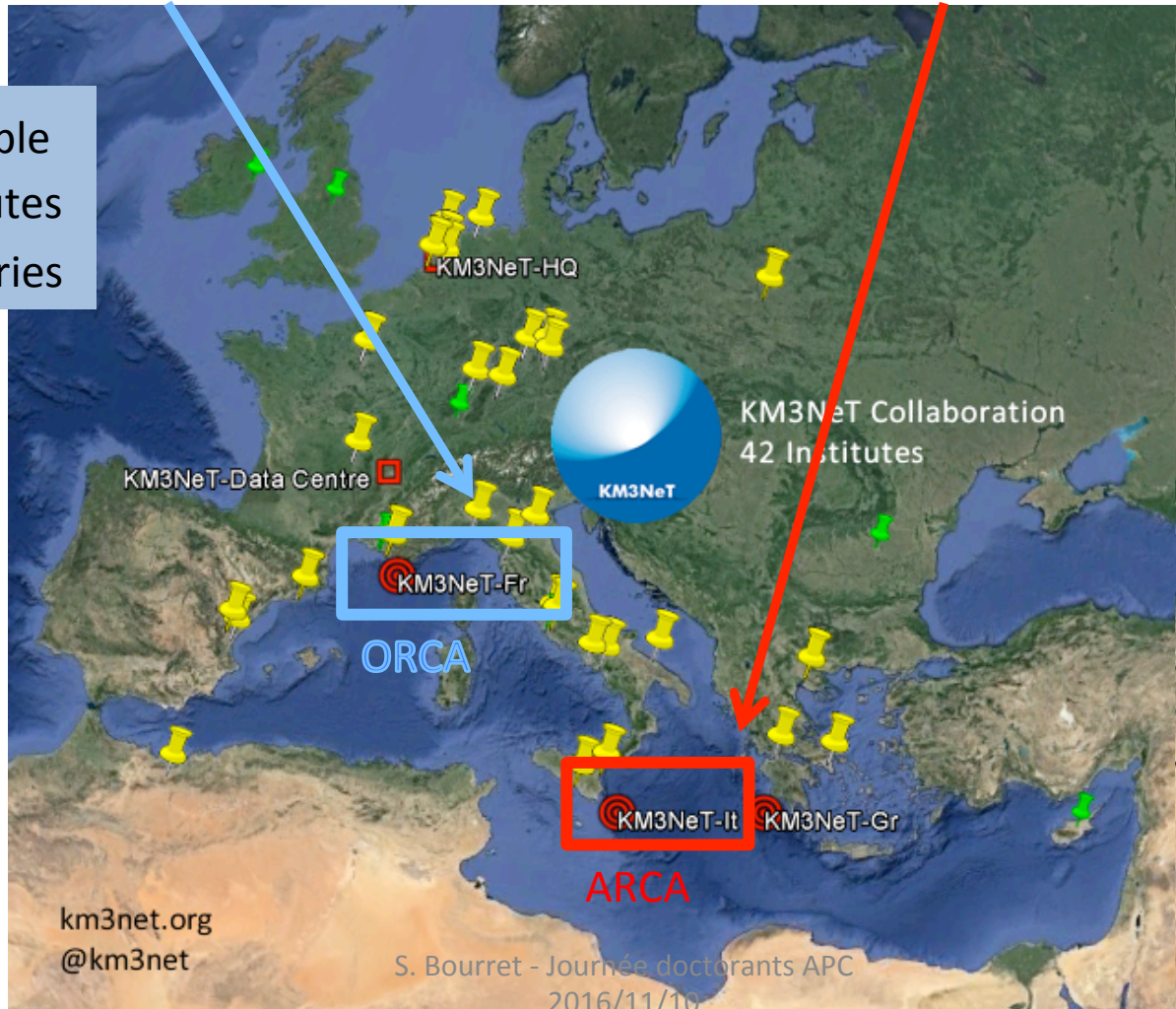
Where & who

KM3NeT is a distributed research infrastructure with 2 main physics topics:

Oscillations and Astroparticle Research with Cosmics in the Abys

Low-Energy studies of atmospheric neutrinos – High-Energy search for cosmic neutrinos

240 people
42 institutes
12 countries



2 sites
currently
under
construction
in France and
Italy:

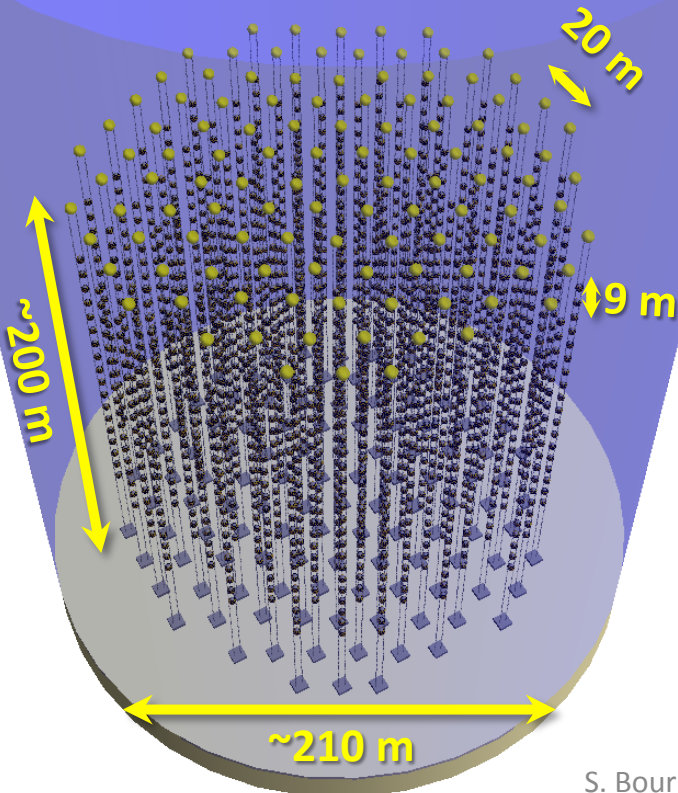
KM3NeT-Fr
(Toulon, close
to ANTARES)

KM3NeT-It
(Capo Passero,
Sicily)



The ORCA detector

- **~5.7 Mt** instrumented
- **115** strings
- **18** DOMs / string
- **31** PMTs / DOM
- Total: **64k*3"** PMTs

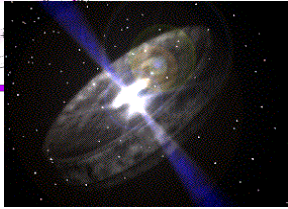
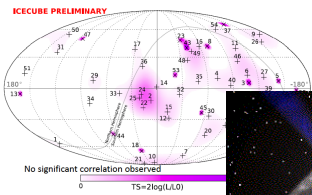


Digital Optical Module



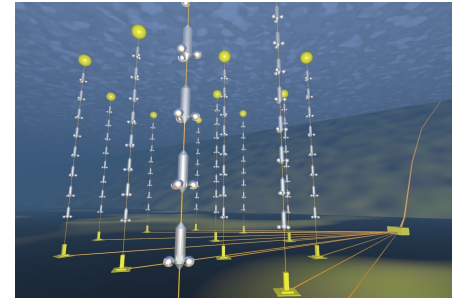
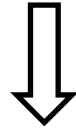
- 31 x 3" PMTs
- Uniform angular coverage
- Directional information
- Digital photon counting
- Background rejection
- All data to shore

Depth=2475m



Antares 2008-2016

Neutrino astronomy



KM3NeT 2015-...

Neutrino astronomy + neutrino oscillations



ARCA

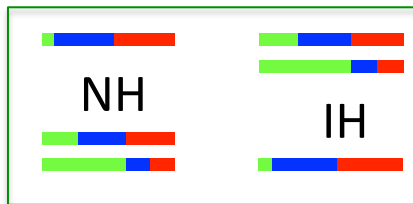
Neutrino astronomy



ORCA

Atmospheric neutrinos

Oscillations



Neutrino Mass Hierarchy
+ other neutrino parameters

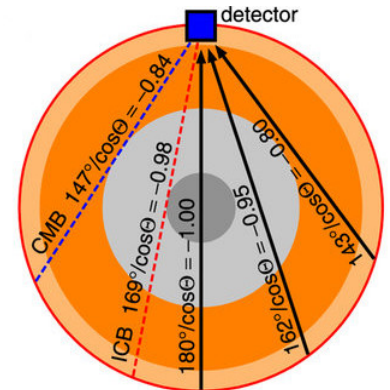
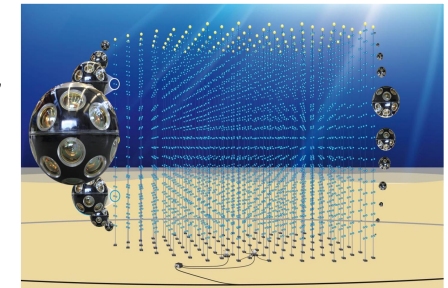
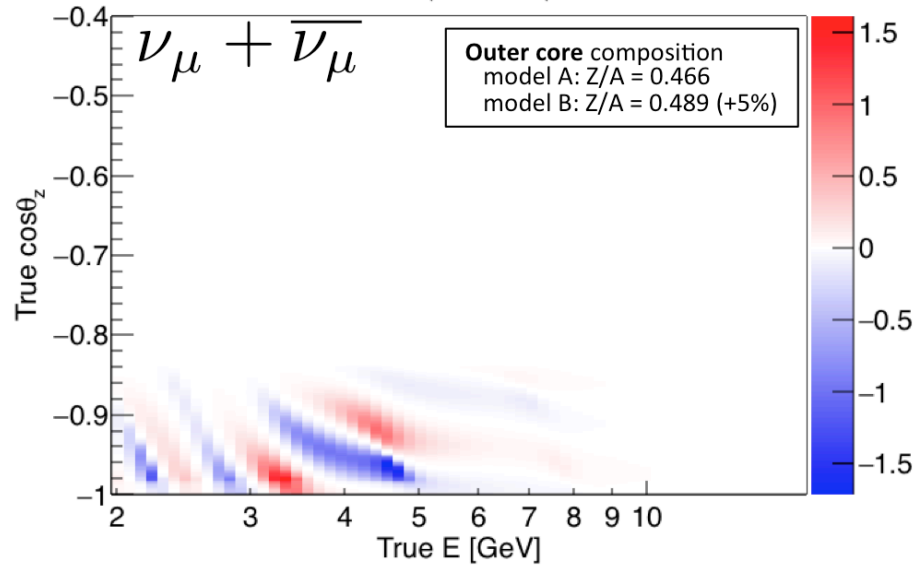


figure C. Rott

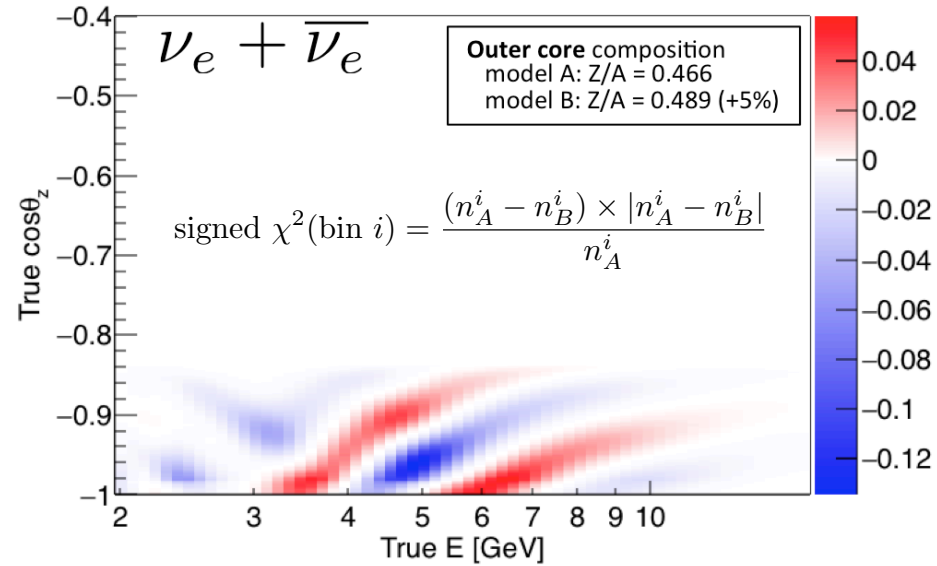
Tomography of the Earth
with neutrino oscillations

Perfect detector signal – Outer core tomography

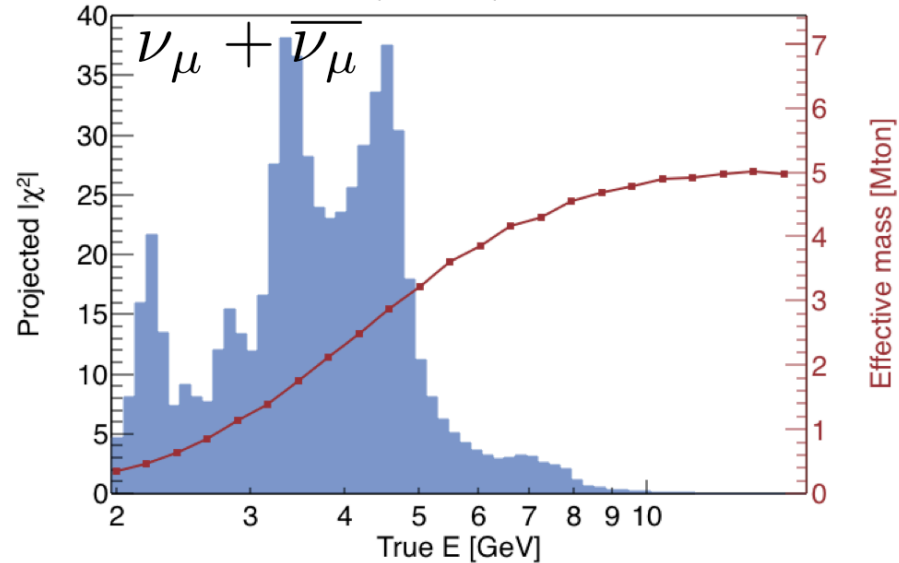
Signed χ^2 for ν_μ CC + $\bar{\nu}_\mu$ CC events



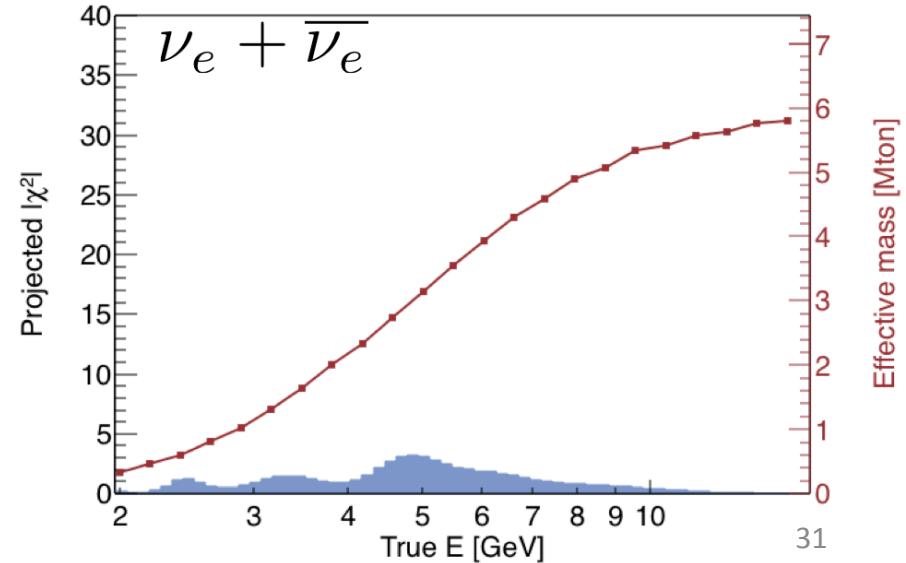
Signed χ^2 for ν_e CC + $\bar{\nu}_e$ CC events



χ^2 for ν_μ CC + $\bar{\nu}_\mu$ CC events

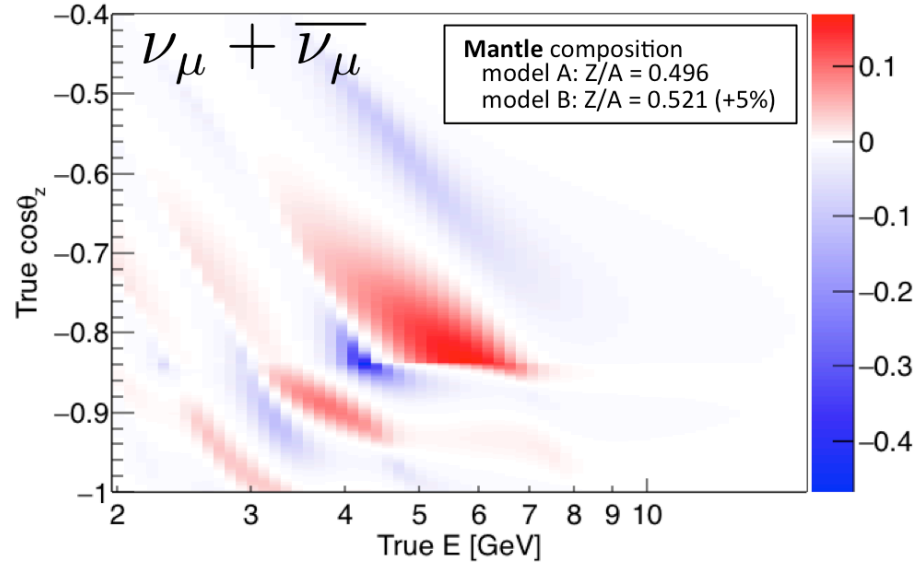


χ^2 for ν_e CC + $\bar{\nu}_e$ CC events

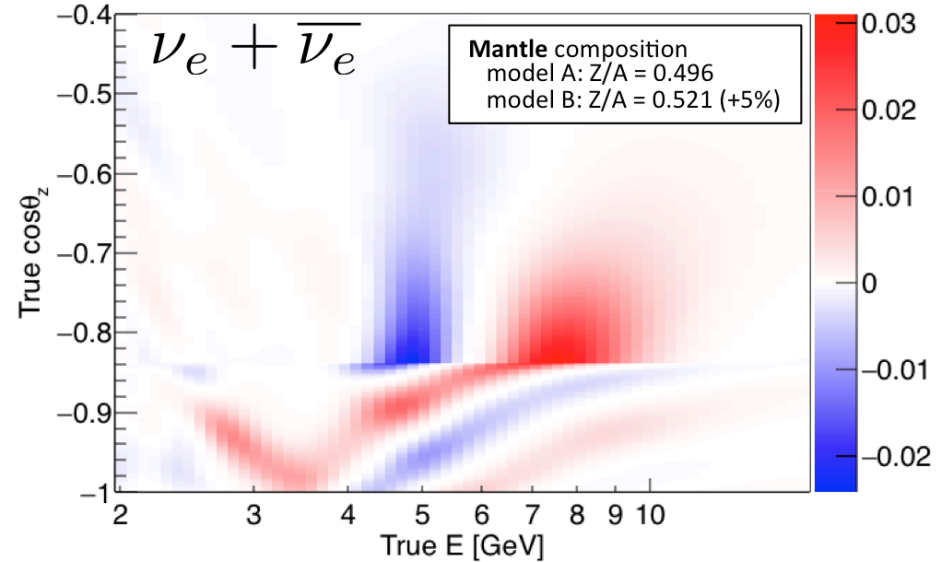


Perfect detector signal – Mantle tomography

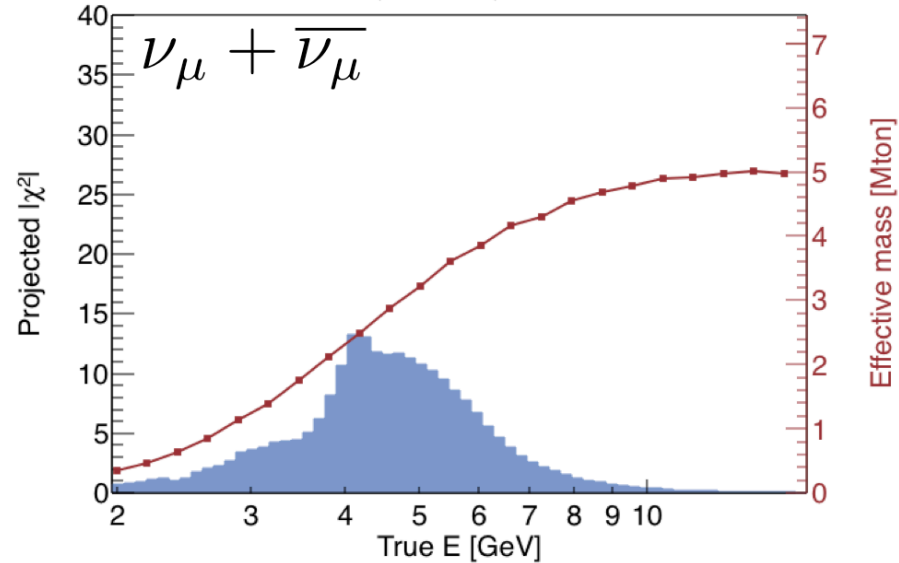
Signed χ^2 for ν_μ CC + $\bar{\nu}_\mu$ CC events



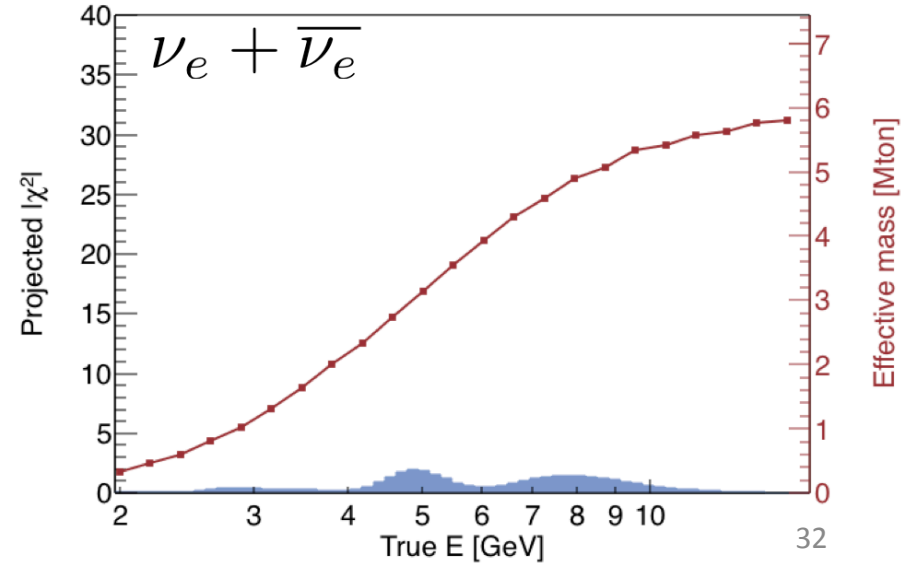
Signed χ^2 for ν_e CC + $\bar{\nu}_e$ CC events



χ^2 for ν_μ CC + $\bar{\nu}_\mu$ CC events



χ^2 for ν_e CC + $\bar{\nu}_e$ CC events



Detector response

- « event-by-event » response stored into 6D sparse histograms
- Discrete channels: 12 for CC events, 4 for NC events
- Atmospheric muon contamination not included yet

- True $(E, \cos \theta_z, Y_{Bj})$
- True channel:
 - CC $\nu_e/\bar{\nu}_e, \nu_\mu/\bar{\nu}_\mu, \nu_\tau/\bar{\nu}_\tau$
 - NC $\nu/\bar{\nu}$

Reconstruction & Classification



- Reco $(E, \cos \theta_z, Y_{Bj})$
- Event classification ("flavour ID"):
 - Track / Cascade / μ_{atm}

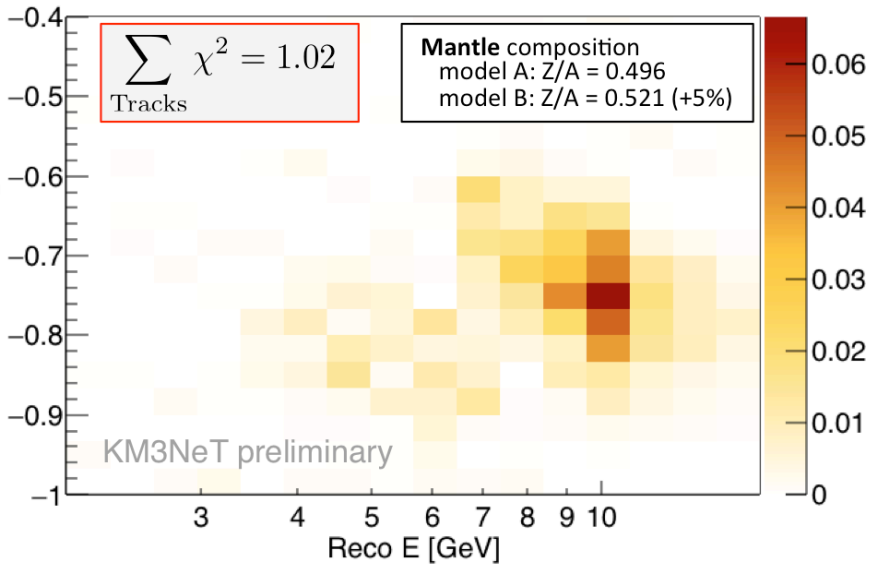
Example of 2D projection of response:

- **Truth** = only one bin
 - $E \approx 10$ GeV
 - $\cos \theta \approx -0.75$
- **Reconstructed** = 2D distribution

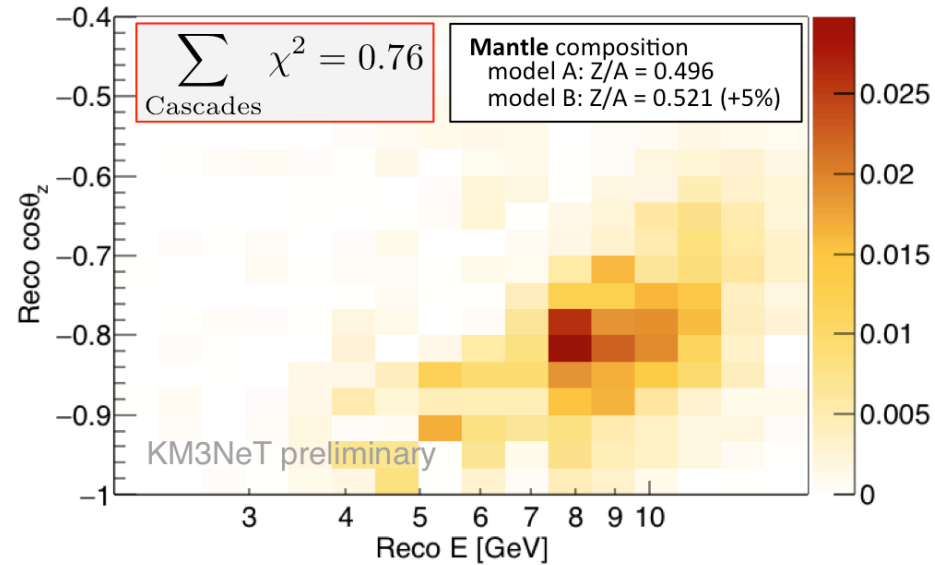


Measured signal: Mantle – ORCA 10 years

2D χ^2 map (E,cos θ) - Tracks

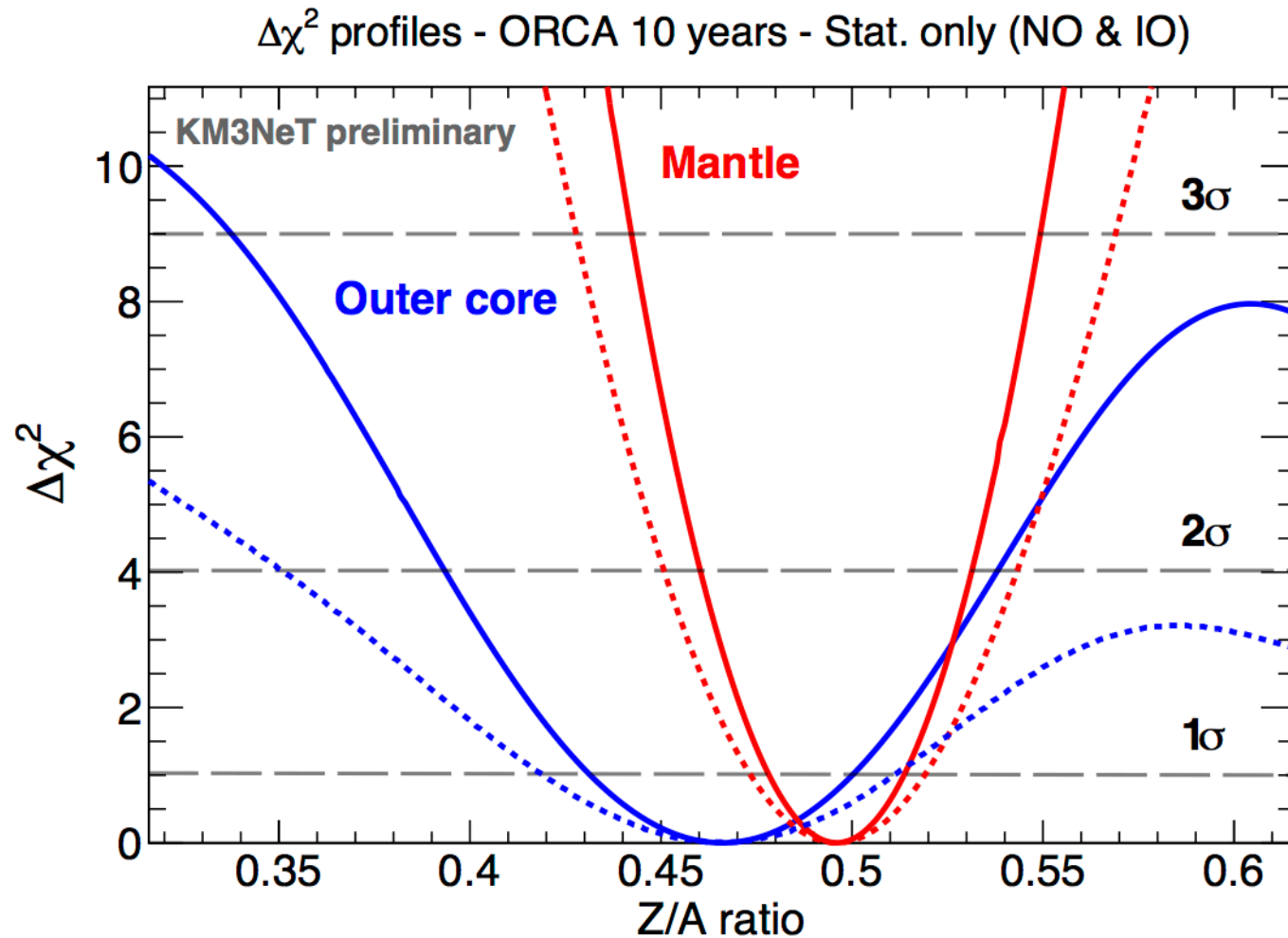


2D χ^2 map (E,cos θ) - Cascades



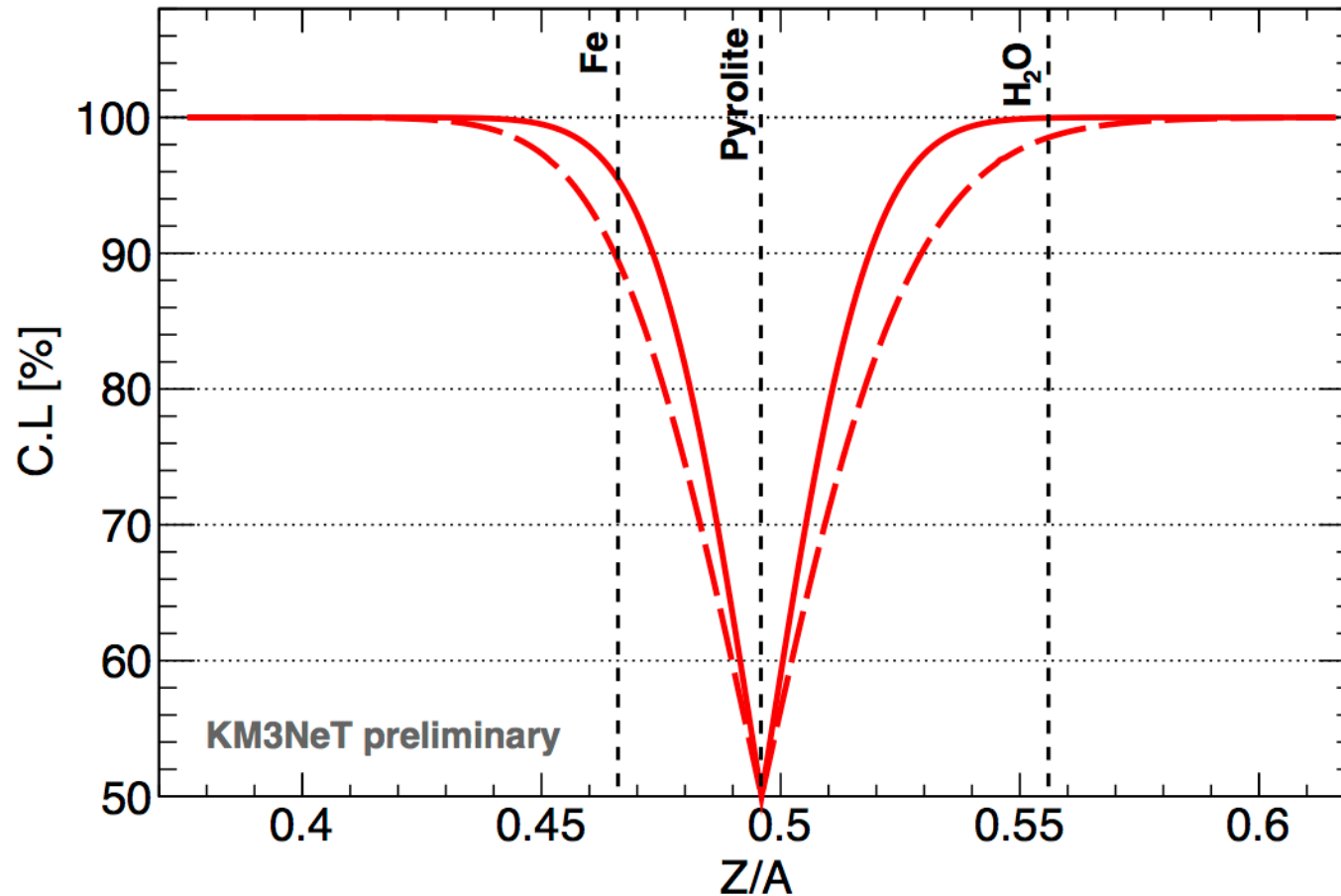
$$\Delta\chi^2 = \sum_{\substack{\text{Tracks,} \\ \text{Cascades}}} \sum_{\substack{\text{bins } E \\ \text{bins } \cos\theta_z}} 2 \left[n_A - n_B + \log \left(\frac{n_B}{n_A} \right) \right]$$

Statistical significance – ORCA 10 years



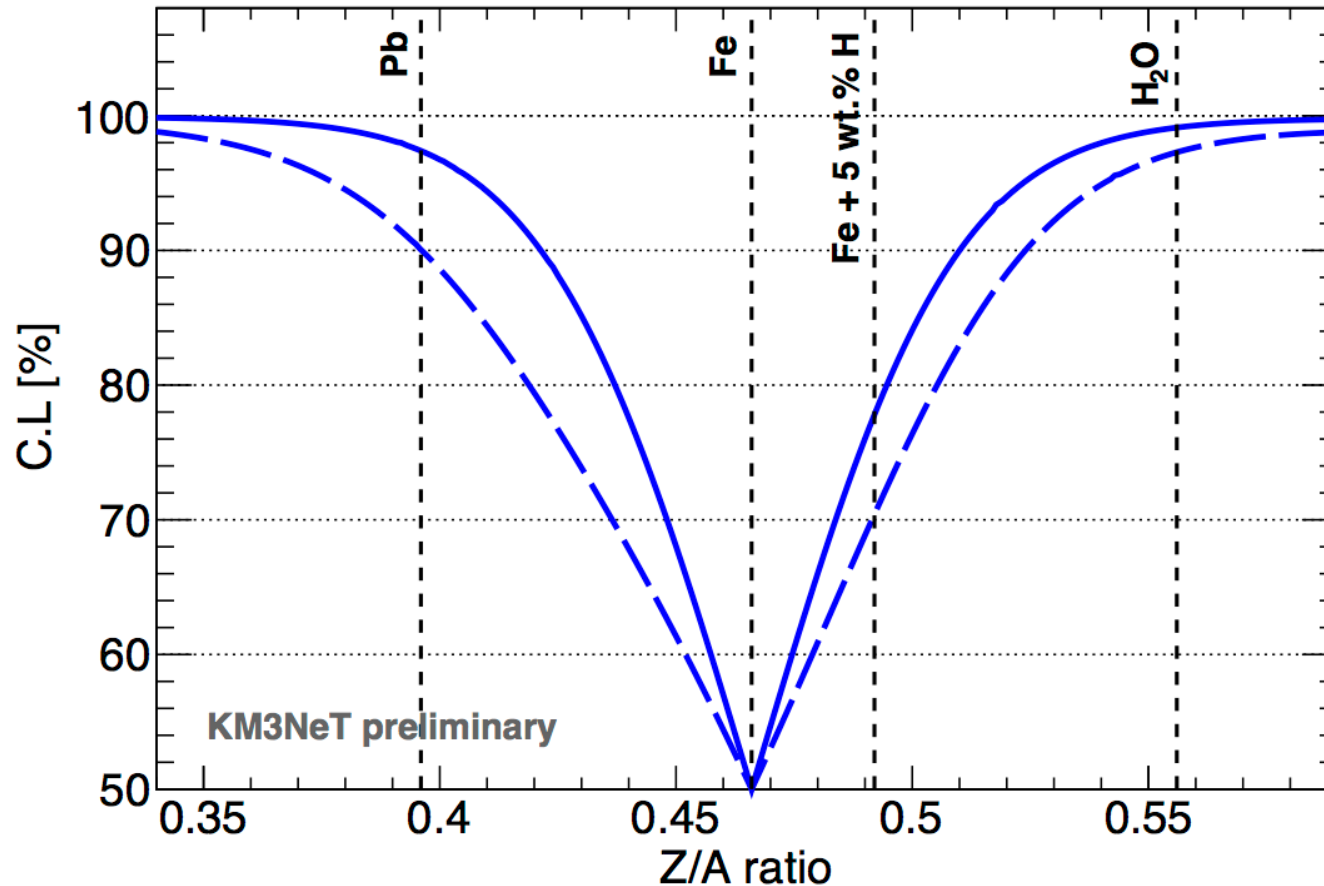
Confidence level for rejecting basic composition of mantle (pyrolite) vs alternative True Z/A

Mantle - ORCA 10 years - Stat. only



Confidence level for rejecting basic composition of outer core (pure Fe) vs alternative True Z/A

Outer core - ORCA 10 years - Stat. only



KM3NeT preliminary